

# ERRATA

## THERMAL RADIATION: AN INTRODUCTION

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**NOTE:** The recent update to the International System of Units has changed some values for constants used in Example Problems and Homework. These are updated in the Errata to use the latest SI values. The recently updated SI constant values are noted at the end of this Errata listing.

### Page Correction

- 37,38 In Examples 2.1 and 2.2, the revised value of  $C_1$  (see page 429, below) changes the results slightly. In Example 2.1, the first equation results in 2747 W/(m<sup>2</sup>·μm·sr), and the final result is 8629 rather than 8627 W/(m<sup>2</sup>·μm). In Example 2.2, replace the result of  $0.256 \times 10^8$  with  $0.259 \times 10^8$ .
- 42 At end of paragraph preceding Eq. (2.28), replace the parenthetical expression with: (assuming  $n$  is independent of wavelength):
- Eq. (2.29) should read\*:  $\lambda_{\max}T = C_3 = 2897.7720 \mu\text{m}\cdot\text{K}$
- 43\* Eq. (2.30), replace the value of  $C_4$  with  $4.095\ 68 \times 10^{-12}$  and in the line below, with  $4.095\ 68 \times 10^{-12}$
- 44\* In Eq. (2.33) replace the value of  $\sigma$  with  $5.670\ 374\ 419 \times 10^{-8}$  (W/m<sup>2</sup>·K) (This affects the value of  $\sigma$  shown in some later examples, but does not affect the results since usually rounded to at most 6 sig figs).
- 47 In Eqs. (2.38) and (2.39), make the dummy variable explicit by using the forms:

$$F_{0 \rightarrow \lambda T} = \frac{2\pi C_1}{\sigma T^4} \int_0^{\lambda} \frac{d\lambda}{\lambda^5 \left( e^{\frac{C_2}{\lambda T}} - 1 \right)} = \frac{2\pi C_1}{\sigma C_2^4} \int_{\zeta}^{\infty} \frac{\zeta^{*3}}{e^{\zeta^*} - 1} d\zeta^* \quad (2.38)$$

$$F_{0 \rightarrow \lambda T} = 1 - \frac{15}{\pi^4} \int_0^{\zeta} \frac{\zeta^{*3}}{e^{\zeta^*} - 1} d\zeta^* \quad (2.39)$$

49 Example 2.9: The result should read = 0.083053 – 0.002134 = 0.0809, giving an efficiency of about 8 %.

63 For clarity, reword the paragraph starting with “We start the analysis...” with:

We start the analysis with the radiative intensity  $I_\lambda(\theta, \phi)$  leaving surface element  $dA$  as in Figure 1.18a. The projected area is formed by taking the area that the energy is leaving and projecting it normal to the direction of the radiation,  $dA \cos \theta$ .

To analyze the radiative exchange between two finite surfaces, we need to carry out integration over the entire area of each surface. For this, consider radiative energy leaving a small area element  $dA_1$  and traveling in a nonparticipating medium as in Figure 1.18b. Assume that this energy is incident on a second small area element  $dA_2$  on finite area  $A_2$ , at distance  $S_{12}$  from  $dA_1$ . The projected areas are formed by taking the area that the energy is passing through and projecting it normal to the direction of the radiation; therefore,  $dA_1 \cos \theta_1$  and  $dA_2 \cos \theta_2$  are the normal components of the infinitesimal areas along direction  $S_{12}$ . The elemental solid angle is centered about the direction of the radiant path and has its origin at  $dA$ . Using the definition of spectral intensity  $I_{\lambda,1}$  as the rate of energy passing through  $dA_1$  per unit projected area per unit solid angle and per unit wavelength interval, the energy  $dQ_{\lambda,1}$  from  $dA$  passing through  $dA_1$  in the direction of  $S_{12}$  is

67 In Problem 2.10, the answer should be 0.809 h

78-80 In Figure 3.4, change the cosine function to  $0.850 \cos(\theta)$ . Use this function in the solution of Example 3.3, and round the result to 3 significant figures to give a final result of 32,200 W/m<sup>2</sup>.

90 In the equation at the top of the page, all upper limits in the integrals should be  $\infty$ , not N (3 places).

91 In Example 3.6, all upper limits in the integrals in the first equation should be  $\infty$ , not N (3 places). Also, due to the revised constants\*, the final answer should read:

$$\alpha = 0.90 \times 0.32639 + 0.37832 + 0.25(1 - 0.88375) = 0.7015.$$

109 In Homework Problem 3.5, the answer to part (a) should be 0.764.

111 The answer to HW Problem 3.10 should be 30.4 min.

128 Example 4.1: Answer for perpendicular component should be 0.299, and reflectivity for unpolarized incident intensity should be 0.251.

132 Table 4.2. in the values for Normal Spectral Reflectivity for Aluminum, from Eq. (4.51), the values should be 0.916 and 0.979 instead of 0.883 and 0.970.

137 Example 4.4: Replace the third equation with:

$$\alpha_n(T = 250 \text{ K}) = \frac{\epsilon_n(T = 500 \text{ K})}{\sqrt{2}} = \frac{\sqrt{1/2} \int_0^\infty \epsilon_{\lambda,n}(T = 500 \text{ K}) I_{\lambda,b}(500 \text{ K}) d\lambda}{\int_0^\infty I_{\lambda,b}(500 \text{ K}) d\lambda}$$

and replace the fourth equation with

$$\begin{aligned} \epsilon_n(T = 500 \text{ K}) &= 0.0348T \sqrt{r_{e,273}} = 0.0348T \sqrt{r_{e,273}} \sqrt{\frac{273}{298}} T \\ &= 0.0348 \sqrt{10 \times 10^{-6}} \sqrt{\frac{273}{298}} \times 500 = 0.053 \end{aligned}$$

147 Last equation: Replace  $q_{\text{sol}}$  with  $G_{\text{sol}}$ .

149 In first line of the final paragraph in Example 4.6, replace 15  $\mu\text{m}$  with 1.5  $\mu\text{m}$

150 Example 4.7: The result of the first equation should be 1150  $\text{W}/\text{m}^2$ . The result of the second equation should be 68.3  $\text{W}/\text{m}^2$ . The final result should be 1150-68 = 1082  $\text{W}/\text{m}^2$ .

151 In HW Problem 4.1, the answers should be 0.9797 and 0.9405.

152 In HW 4.6, the answers should be 645 K; 0.0329; 1.31

In HW 4.8, The answers should be: *Answer:*  $\epsilon_{\lambda,Al}(\lambda = 0.484 \mu\text{m}) = \underline{0.086}$  ;  
 $\epsilon_{\lambda,Al}(\lambda = 8.06 \mu\text{m}) = \underline{0.014}$ ;  $\epsilon_{\lambda,Ti}(\lambda = 0.484 \mu\text{m}) = \underline{0.678}$ ;  $\epsilon_{\lambda,Ti}(\lambda = 8.06 \mu\text{m}) = \underline{0.043}$ .

In HW 4.9, add "...when exposed to the sun." to the problem statement.

428\* The speed of light,  $c_0$ , should read  $c_0 = 2.997\,924\,58 \times 10^8 \text{ m/s}$

The reduced Planck's constant should read  $\hbar = h/2\pi = 1.054\,571\,817 \times 10^{-34} \text{ J}\cdot\text{s}$

The value for the Boltzmann constant should read  $k = 1.380\,649 \times 10^{-23} \text{ J/K}$

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The value for the classical electron radius should read:

$$2.817\,940\,3262(\pm 13) \times 10^{-15} \text{ m}$$

The value for the electron volt should read:  $1 \text{ eV} = 1.602\,176\,634 \times 10^{-19} \text{ J}$

429\*

The value of the radiation constant  $C_1$  in SI should read

$$C_1 = 0.595\,521\,49 \times 10^8 \text{ W}\cdot\mu\text{m}^4/(\text{m}^2\cdot\text{sr})$$

$$C_1 = 0.595\,521\,49 \times 10^{-16} \text{ W}\cdot\text{m}^2/\text{s}$$

The value of the Stefan-Boltzmann constant in SI should read

$$\sigma = 5.670\,374\,419\dots \times 10^{-8} \text{ (W/m}^2\cdot\text{K)}$$

*\*Most of these changes result from a recent fundamental revision in the International System of Units. See Appendix N in the on-line Appendices at:*

<http://www.thermalradiation.net/appendix.html>

*or the video at*

<http://www.thermalradiation.net/videos.html>